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Aspect sensitivity of VHF radar echoes observed in the middle and upper troposphere during the passage of a cut-off low

Bernard Campistron

Laboratoire d'Aérodologie (CNRS), Observatoire Midi-Pyrénées, Université Paul Sabatier, Toulouse, France

Yves B. Pointin

Laboratoire de Météorologie Physique (CNRS), Observatoire de Physique du Globe de Clermont-Ferrand
Université Blaise Pascal, Clermont-Ferrand, France

Fabienne Lohou and Jean-Pierre Pagès

Laboratoire d'Aérodologie (CNRS), Observatoire Midi-Pyrénées, Université Paul Sabatier, Toulouse, France

Abstract. The aspect sensitivity of VHF radar signals is related to the enhancement of the echo power and the narrowing of the spectral width at vertical incidence relative to the off-zenith directions. The strong aspect sensitivity is usually associated with stable atmospheric layers. In this paper the aspect sensitivity measured between an altitude of 4 and 15 km during the passage of a tropopause folding linked to a cut-off low is analyzed and discussed. The study mainly uses data acquired during a 2-day period with a three-beam VHF radar with relatively broad beam widths (6.8°). The angular dependence of the echoes is investigated using the power ratio of vertical to oblique beam and the spectral width obtained at vertical incidence. These two diagnostic parameters provided a nearly identical depiction of the stratospheric descent associated with the upper frontal zone. However, a discrepancy is observed between the location of the stratospheric subsidence and the tropopause height offered by the radar and that deduced from the meteorological analysis produced by an operational forecasting numerical model. Owing to the aspect sensitivity of the echoes, an underestimate amounting to 20% was found in the horizontal wind measurement. Finally, a comparison of simultaneous observations of two closely spaced VHF radars (100 km) has shown a good spatial coherence of the aspect sensitivity structures.

1. Introduction

The so-called aspect sensitivity of VHF radar measurements is related to the zenith angle dependence of the backscattered power and spectral width. The enhancement of the echo power and the narrowing of the spectral width at vertical or near-vertical incidence relative to the off-zenith directions are of common occurrence and well established now for VHF radar observations [e.g., *Gage and Green*,

1978; *Röttger et al.*, 1981; *Hocking et al.*, 1986; *Woodman and Chu*, 1989; *Yoe et al.*, 1994]. Strong aspect sensitivity is mainly found in the stratosphere and mesosphere. The unstable and turbulent troposphere usually presents slight aspect sensitivity except in the stable region around the tropopause, which exhibits the strongest vertical enhancement in the lower atmosphere. This was used as a mean for the tropopause height detection [*Gage and Green*, 1978]. Occasionally, aspect sensitivity enhancement was observed in the middle and lower troposphere, particularly during the passage of frontal zones [e.g., *Hooper and Thomas*, 1995].

There is still controversy about the backscattering process giving rise to the echo angular dependence,

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although it is generally accepted that it is produced by stable and thin horizontally stratified atmospheric structures. The existence of thin temperature sheets and their relation with VHF signal enhancement was proved by *Luce et al.* [1995] using very high resolution in situ measurements. Two main echo mechanisms have been proposed to explain the origin of the aspect sensitivity of the echoes. (1) The first is scattering from anisotropic turbulence [e.g., *Gage and Green*, 1978; *Doviak and Zrnic*, 1984; *Hocking and Hamza*, 1997]. *Woodman and Chu* [1989] have proposed a model of a thin atmospheric sheet in which the scattering sources evolve from anisotropic structures at the edges toward more isotropic ones when approaching the center of the layer. (2) The second is partial reflection from a single or a group of thin horizontal layers with sharp vertical variations in the refractive index. The specular echo model (mechanism 2) is called Fresnel scattering when the layers are vertically randomly distributed within the radar resolution volume. It is called diffuse reflection when the layers are not perfectly horizontal but present some roughness due to waves or turbulence [e.g., *Röttger and Liu*, 1978; *Röttger*, 1980; *Gage et al.*, 1981]. According to *Hocking* [1997], the distinction between echoes from anisotropic turbulence and from an ensemble of specular reflectors is always difficult in radar measurements. In the following, we will use the general term of anisotropic specular reflection for these two types of echoes.

Apart from the backscattering from hydrometeors, scattering from refractive index irregularities induced by isotropic turbulence [*Tatarskii*, 1961] is a well-accepted mechanism [*Woodman and Chu*, 1989] to explain radar echoes observed in the absence of aspect sensitivity, which usually predominates for zenith angles larger than about 10° [*Hocking et al.*, 1990]. Certainly, it is the conjunction of these different kinds of echoing sources, more or less dominant within the radar resolution volume, which produces the observed radar signal.

In this study, we present an analysis of the aspect sensitivity observed during the passage of a jet stream linked to a cut-off low. In particular, we show that aspect sensitivity might be helpful to map stratospheric descent not well resolved by a forecasting model.

2. The Experimental Setup

The observations discussed here have been collected during the ESTIME (Echanges Stratosphère Troposphère Investigations à Moyenne Echelle) experiment (1993-1995), which was aimed at investigating the stratosphere-troposphere exchanges and mixings associated with cut-off low events. The project was mainly based on data acquired by the regional wind-profiler network (5-VHF and 1-UHF radars) deployed in southern France, roughly arranged in a triangular formation with a mean side dimension of 400 km [*Campistron*, 1997]. The present study focuses on the data of one of the VHF radars situated at the Centre de Recherches Atmosphériques (CRA) on a plateau (43.13°N , 0.37°E) at a height of 600 m in the foothills of the Pyrénées mountains (see Figure 1). This mountain range, separating France and Spain between the Atlantic Ocean and the Mediterranean Sea, is oriented roughly west to east. It is an isolated structure with a length of about 400 km and a mean transept of 50 km. The nearest summit at a distance of 30 km from the radar culminates at 2800 m. In section 7, the analysis of the CRA radar data is complemented with a comparison made with observations collected, in a restricted period of 10 hours, by another VHF radar of the network of identical instrumental characteristics. This radar, located in Toulouse (43.62°N , 1.43°E) at an altitude of 150 m, is approximately 100 km northeast of the CRA radar (see Figure 1).

These two radars were operated at a 45-MHz transmitted frequency, with a 6-kW peak power, a 6400-Hz pulse repetition frequency, and a 25- μs -coded pulse. A 10-element complementary code provided an effective range resolution of 375 m. A vertical beam and two oblique beams, with perpendicular azimuth angles and an off-zenith direction of 14.7° , were alternately used to get vertical profiles of the three components of the wind every 12 min. A coaxial-collinear antenna ($56 \times 56 \text{ m}^2$) provided for these three directions a one-way half-power beam width θ_b of 6.8° . The vertical sampling was made of 48 range gates, evenly spaced by 362 m, starting at an altitude of 3839 m.

The measurements were off-line carefully edited in order to extract the three moments of the meteorological spectra. In a first step, the mean noise level of the Doppler spectra was determined with the

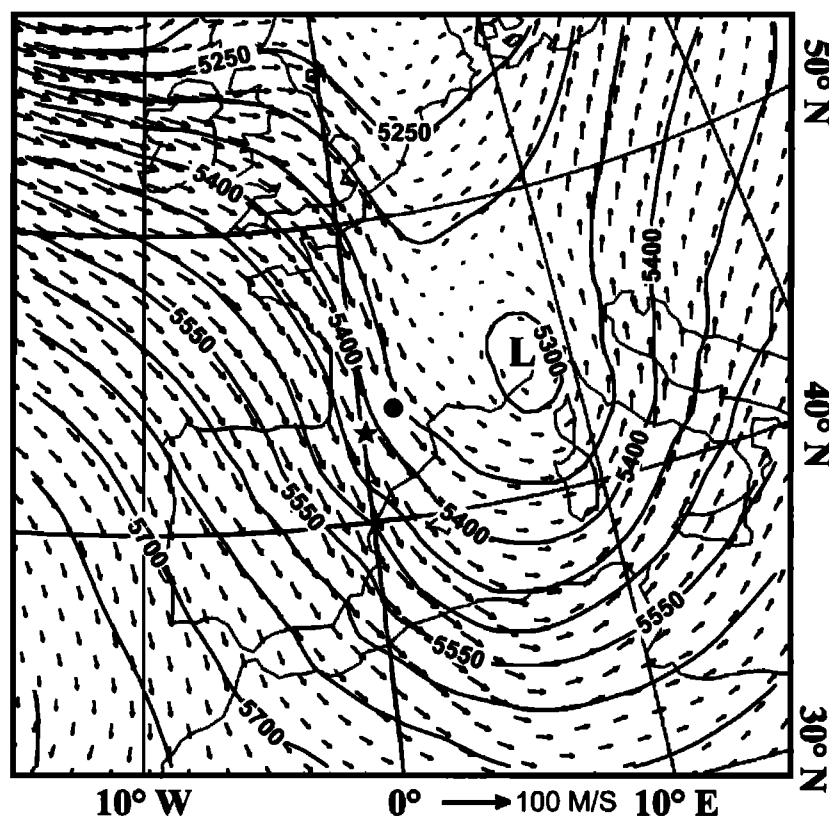


Figure 1. Horizontal wind and height contours (meters) on the 500-hPa surface at 1200 UTC on March 4, 1995, provided by the European Centre for Medium-Range Weather Forecasting (ECMWF) model. The star and the circle, nearby the 0° meridian in the south of France, give respectively the location of the Centre de Recherches Atmosphériques (CRA) and Toulouse (TLSE) VHF radars.

objective method proposed by *Hildebrand and Sekhon* [1974], and the zero velocity spectral line, contaminated by ground clutter and instrumental effects, was removed and replaced by a value interpolated from the two adjacent spectral lines. Usually, several distinct Doppler peaks emerge above the noise level. In a second step, the selection of the meteorological peak was done on the basis of vertical and temporal continuity and thresholdings. Finally, the selected peak was fitted to a Gaussian curve from which mean radial velocity, echo power, and spectral width were computed.

3. Meteorological Overview of the Case Study

Early on March 3, 1995, a baroclinic wave of the westerly jet stream begins to develop over western

Europe and the nearby Atlantic Ocean. Figure 1 displays, on the 500-hPa surface, the meteorological analysis of the European Center for Medium-Range Weather Forecasts (ECMWF) model obtained on March 4, 1995, at 1200 UTC. Figure 1 shows a cold air outbreak in a strong northerly flow crossing the British Isles, the west of France, and the north of Spain. A deepening low, centered above the Alpine Mediterranean coast, is located on the trough extending between the North Sea and Corsica. On March 5 at 1200 UTC, the low has moved toward central Italy and is now almost detached from the synoptic current. The day after, this cut-off low has disappeared from the meteorological charts.

The data analysis is restricted to the period from 1600 UTC March 3 to 0900 UTC March 5, 1995. The time-height section of the wind observed by the radar is presented in Figure 2a. As for the next radar

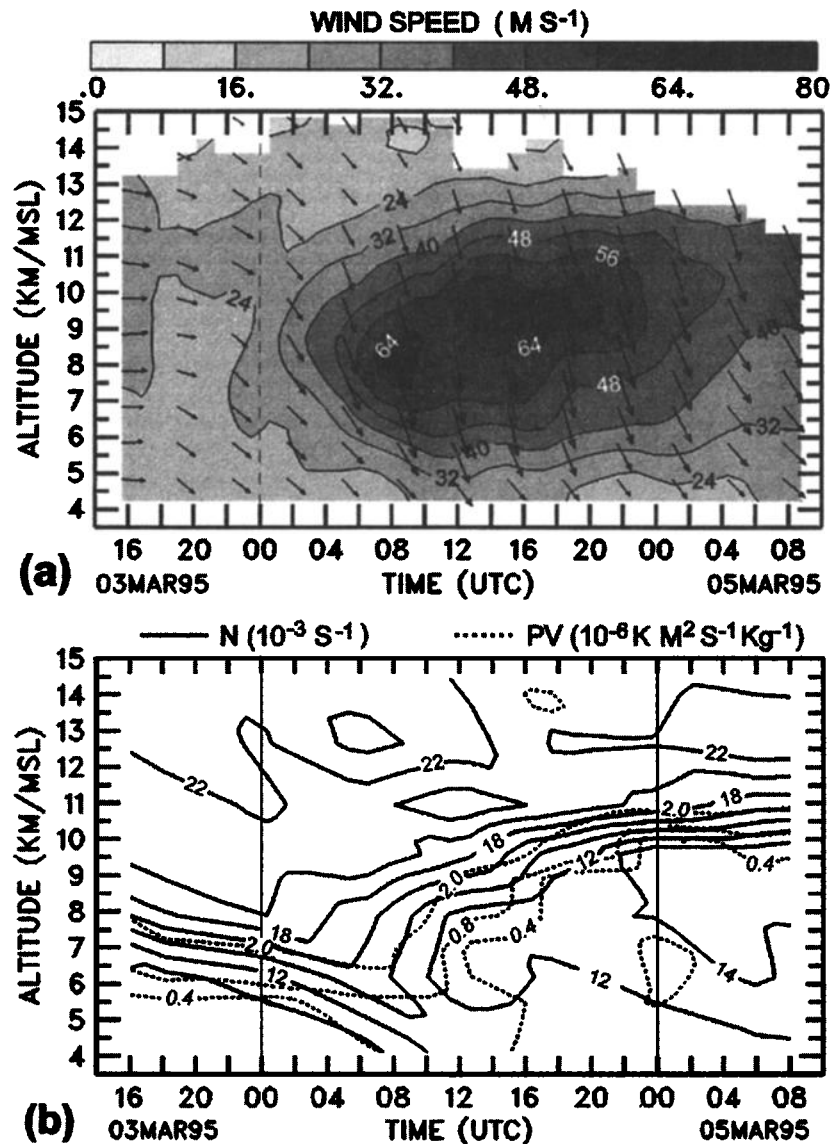


Figure 2. Time-height sections above the radar site: (a) horizontal wind vectors and speed (m s^{-1}) observed by the VHF radar, and (b) contours of Brunt-Väisälä frequency N (solid lines) and potential vorticity (PV) (dashed lines) derived from the analysis of the ECMWF model.

sections, a smoothing in height (three data gates) and in time (three hours) was applied to the data for better clarity. The southward descending branch of the jet streak associated with the baroclinic wave appears on March 4, between 0600 and 2000 UTC, as a core of strong wind centered at an altitude of about 9 km, reaching values larger than 65 m s^{-1} around 1600 UTC.

The time-height section above the radar site of the Brunt-Väisälä frequency N and potential vorticity (PV) are shown in Figure 2b. These fields were obtained from an analysis of the ECMWF model. N is a measure of the static stability of the atmosphere, and PV is a conservative quantity for inviscid adiabatic dry motion. These parameters are expressed by the following equations:

$$N = \frac{g}{\theta} \frac{\partial \theta}{\partial z}, \quad PV = \xi_a \frac{\partial \theta}{\partial p}$$

where g is the gravitational constant, θ is the potential temperature, ξ_a is the vertical component of the absolute vorticity, z is the altitude, and p is the pressure.

The packing of the N contours indicates a tropopause height of about 7 km at the beginning of the period and of about 10.5 km at the end of the period. On March 4 between 0400 and 1400 UTC, the downward sloping of the contours reveals the passage of the baroclinic zone linked to an upper level frontal surface. A similar feature is observed on the PV contours and corroborates the intrusion of stratospheric air into the upper troposphere. According to the 2- PV -unit contour, which is usually taken as a separation between stratospheric and tropospheric air, that stratospheric descent, which precedes the passage of the jet core, extends down to an altitude of about 6.5 km on March 4. In view of the previous synoptic description, this weak tropopause fold is situated on the cyclonic side of the jet stream.

4. Aspect Sensitivity Analysis

This study of aspect sensitivity is based on three different approaches, often reported for the tropopause height detection. The first method analyzes only the local enhancement of the echo power P_v obtained at vertical incidence [Gage and Green, 1978]. The second one makes use of the P_v/P_o ratio between the echo power observed at vertical and that observed in an oblique direction [Tsuda *et al.*, 1986], where P_o represents the average of the echo power received by the two oblique beams. The last method is based on the spectral width σ_v measured in the zenith direction [Hocking *et al.*, 1986]. The spectral width is here defined as the root-mean-square of the Doppler spectrum variance. For a vertically pointing beam with a relatively large aperture, the beam-broadening effect governs mostly the spectral width. According to Sloss and Atlas [1968], a finite beam induces a spectral width σ_b , which is expressed for isotropic scatters and for a Gaussian beam diagram by the following relationship:

$$\sigma_b = \theta_b (16 \ln 2)^{-0.5} V \quad (1)$$

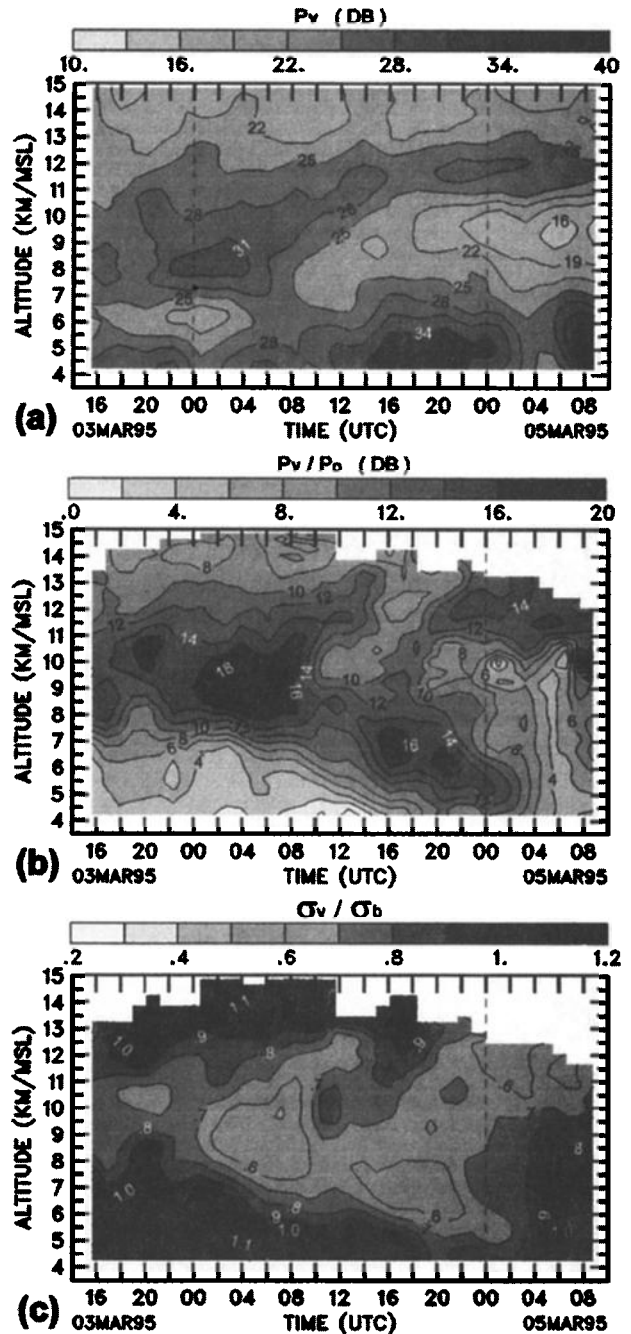


Figure 3. Time-height sections of aspect sensitivity parameters observed by the VHF radar: (a) vertical echo power P_v (decibels relative to an arbitrary reference power level); (b) vertical to oblique echo power ratio P_v/P_o (decibels); and (c) normalized vertical spectral width σ_v/σ_b .

where V is the horizontal wind speed (m s^{-1}), and θ_b (radians) is the half-power one-way width of the beam.

When the beam broadening is the only source of spectral broadening, then isotropic scatterers give σ_v value equal to σ_b , and anisotropic specular reflectors give σ_v value smaller than σ_b [Hocking *et al.*, 1990]. If other spectral broadening factors, such as turbulence, add to the beam effect, σ_v becomes larger than σ_b for isotropic scatterers, but it can become smaller or larger than σ_b for anisotropic specular reflectors depending on the contribution of the extra broadening factors. Therefore the ratios σ_v/σ_b (as P_v/P_o) provide information on the degree of aspect sensitivity of the echoes.

Figure 3a presents the time-height section, for the considered period, of the uncalibrated vertical signal-to-noise ratio P_v corrected for the squared range dependence. In the upper part of the tropopause fold, evidenced by the N field in Figure 2b, there is a core of P_v larger than 30 dB extending slightly downward and centered at a height of 8.5 km on March 4 at 0200 UTC. This core is part of a tongue of P_v enhancement, sloping upward with time, reaching a height of 12 km on March 5. This particular zone, well detached from the environment at the end of the period, follows the tropopause region deduced from Figure 2b. In fact, the so-deduced radar tropopause appears shifted upward from the ECMWF tropopause of approximately 1.5 km.

Time-height sections of P_v/P_o and σ_v/σ_b are presented in Figures 3b and 3c, respectively. It is clearly apparent that these two fields depict a similar pattern of the aspect sensitivity, with opposite variations such that strong values of P_v/P_o are well correlated to small σ_v/σ_b values. However, the P_v/P_o presentation displays much more detail on the aspect sensitivity (and therefore on the stability of atmosphere) than the σ_v/σ_b field. Figure 3c also shows that less than 10% of the data have σ_v/σ_b larger or equal to 1, which is a necessary (but not sufficient) condition for the presence of isotropic echoes. These data are on the average associated in Figure 3b with P_v/P_o smaller than 6 dB, which indicates that come from anisotropic specular reflections. It appears that for this case study, anisotropic specular reflections are preponderant in the middle and upper troposphere.

These aspect sensitivity fields present notable similarities with the vertical echo power field. Corresponding with the core of strong of P_v enhancement in the region of the upper tropopause fold, there is a core of strong aspect sensitivity with values reaching 20 dB for P_v/P_o and less than 0.5 for σ_v/σ_b . In the same way, the upper tropopause is marked by an ascending branch of high aspect sensitivity, which is crossed, on the afternoon of March 4, by a zone of weaker aspect sensitivity connected with the jet stream maximum. This clearly apparent region might indicate different evolutions (mixing, radiative or convective warming or cooling, and so on) than its environment.

The more striking feature of Figures 3b and 3c is a tongue of high aspect sensitivity sloping downward with time, from the upper tropopause fold down to a height of 5 km at the end of March 4. No equivalent feature is evident in the P_v field of Figure 3a, in the N , or in the PV field in Figures 2a and 2b. Probably, this stable zone displayed by this feature of the aspect sensitivity fields is linked to the upper frontal surface not well resolved by the ECMWF model and possibly corresponds to a stratospheric descent.

5. Aspect Angle and Wind Measurement Errors

Aspect sensitivity induces a deformation of the polar diagram of the radar beam which is more strongly weighted toward zenith. As a consequence, the effective pointing angle θ_e of an oblique beam is displaced toward zenith. In order to avoid this source of error in the retrieval of the horizontal wind velocity, θ_e has to be used in place of the nominal value θ_T . The estimation of θ_e requires knowledge of the polar diagram of the scatterers. Hocking *et al.* [1990] proposed a simple model, used here, in which the scatterers have a polar diagram of the form $\exp(-\theta^2/\theta_s^2)$, where θ is the tilt angle of the beam and θ_s is the aspect angle corresponding to the e^{-1} half width of the polar diagram of the backscatterers. This parameterization is in accordance with the Gaussian-like angular dependence of the backscattered power (at least for angles close to the vertical) reported by numerous authors [Röttger *et al.*, 1981; Waterman *et al.*, 1985; Tsuda *et al.*, 1986; Chu *et al.*, 1990]. The

parameter θ_s is related to the degree of aspect sensitivity of the echoes and ranges from 0° for pure anisotropic specular reflections to infinity for pure isotropic scatterers. As shown by *Hocking et al.* [1990], the importance of θ_s relies on the fact that its inverse is directly proportional to the horizontal correlation length of the scattering field.

The estimation, from the radar data, of the effective pointing angle θ_e and of the aspect angle θ_s is deduced using the hypothesis that the scatterers have a backscatter polar diagram of a Gaussian type [*Hocking et al.*, 1990]. It is also assumed that θ_s is independent of the radar beam tilt angle or, in other words, that only one type of scattering process exists. The three following equations are derived by *Hocking et al.* [1990]:

$$\sin \theta_e = \sin \theta_T (1 + \theta_o^2 / \theta_s^2)^{-1} \quad (2)$$

where $\theta_o = \theta_b (8 \ln 2)^{-0.5}$ is the e^{-1} half width of the assumed-Gaussian radar polar diagram.

An estimation of the aspect angle θ_s can be deduced from the powers P_v and P_o obtained in the vertical and oblique directions:

$$\theta_s^2 = \theta_T^2 / \ln (P_v / P_o) - \theta_o^2 \quad (3)$$

Another measure of θ_s can be derived from the spectral width σ_v of the zenith beam Doppler spectra and σ_b the isotropic spectral width defined by (1) from beam broadening, as

$$\theta_s^2 = \theta_o^2 ((\sigma_v / \sigma_b)^2 - 1)^{-1} \quad (4)$$

In (4), the assumption is also made that the spectral width σ_v only results from the beam broadening effect. When this condition is verified, the ratio σ_v / σ_b reaches a maximum value of 1 for isotropic scattering, and the aspect angle θ_s tends to infinity. Actually, other spectral broadening factors, such as turbulence, can add to σ_v and give a σ_v / σ_b value greater than 1. In that case, (4) is no more appropriate and the value of θ_s is arbitrarily set to a large number. Figure 3c shows that this occurred for less than 10% of the analyzed data.

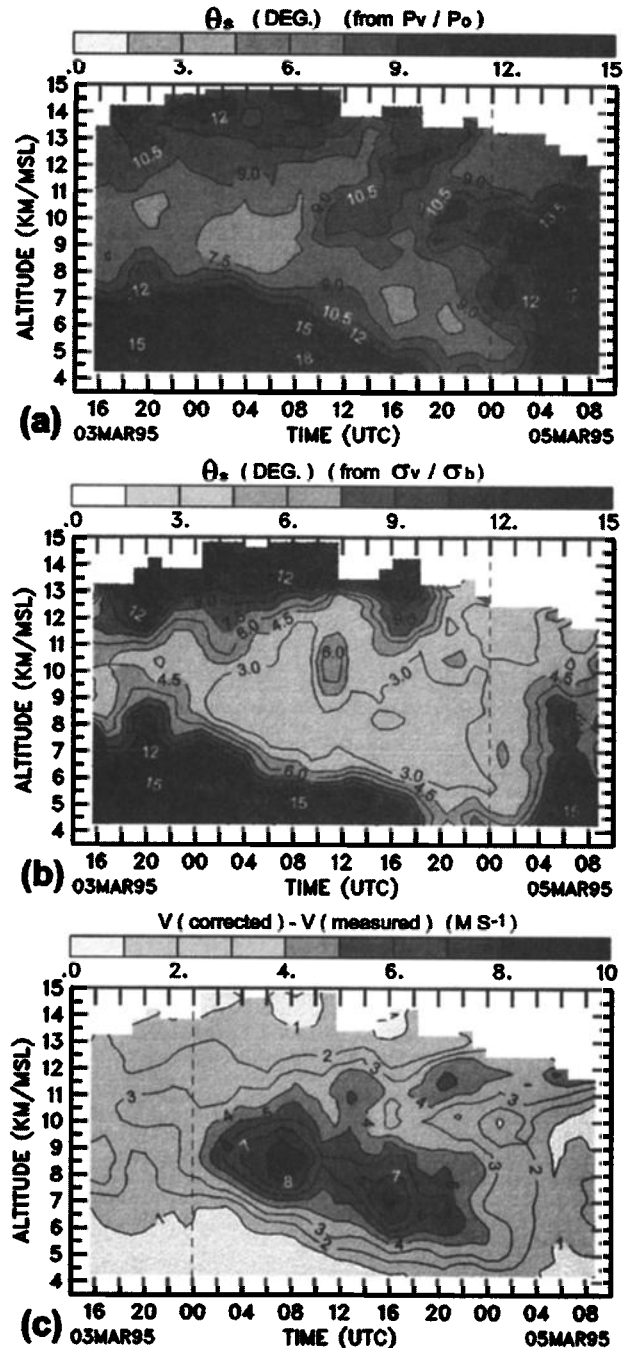


Figure 4. Time-height sections of aspect angle θ_s deduced (a) from P_v/P_o , (b) from σ_v/σ_b , and (c) the difference between the measured wind speed and the corrected wind speed taking into account the effect of the aspect sensitivity.

Time-height sections of the aspect angle derived from the power ratio (equation (3)) and from the vertical spectral width (equation (4)) are presented in Figures 4a and 4b, respectively. For the clarity of the presentation, θ_s values greater than 20° are assigned to 20° . The morphology of the aspect sensitivity given by the aspect angle fields is, as expected, similar to that displayed by the aspect ratio sections of Figures 3b and 3c. On the basis of these figures, the transition from anisotropic specular reflections to more isotropic echoes takes place for θ_s angles exceeding 10° – 15° . This threshold was also noted by Yoe *et al.* [1994] and Hooper and Thomas [1995].

Actually, in Figures 4a and 4b, the values of θ_s making use of the oblique beam power (equation (3)) are always larger than that based on the spectral width measured at vertical incidence (equation (4)). For instance, in the region of large aspect sensitivity, at a height of 9 km at the beginning of March 4, θ_s deduced from (3) reaches a value of about 7° , whereas θ_s deduced from (4) shows value as low as 2° . A similar strong discrepancy between both θ_s estimates was also observed by Hocking *et al.* [1990]. Using a combination of different oblique beams, these authors, Hooper and Thomas [1995], and Jain *et al.* [1997] found a general increase of θ_s with increasing radar observation zenith angle. The zenith angle dependence of the aspect angle is explained by Hocking *et al.* [1990] by the simultaneous presence in the radar resolution volume of different sources of echoes ranging from anisotropic specular structures, more sensitive to vertical beam, to isotropic scatterers that dominate the signal of oblique beams with pointing angles larger than about 10° . The angular dependence of θ_s , in contradiction with the assumption made in the derivation of equations (3) and (4), can explain a large part of the discrepancies between the two methods of θ_s evaluation.

Once the aspect angle is determined, the effective zenith angle θ_e computed from (2) allows us to take into account the aspect sensitivity effect in the retrieval of the horizontal wind velocity. In spite of the limitations of the backscattering model, Yoe *et al.* [1994] found, in comparison with rawinsoundings, a real improvement of the wind velocity estimation when θ_e was used in place of the nominal value θ_T . Note that such corrected wind speed has been used in Figure 2a and in (1).

The time-height section of the difference between the aspect corrected and uncorrected wind speed is presented in Figure 4c. The aspect angle derived from (3) was used instead of that from (4) since it makes use of oblique and vertical information and seems more representative of a mean θ_s than the value extracted from (4). On the other hand, this choice is also dictated by the satisfactory wind correction obtained by Yoe *et al.* [1994] using the aspect angle derived from the power ratio method. The difference between both wind estimates reaches a maximum of 10 m s^{-1} in a zone located in advance and beneath the jet stream core. This corresponds to an underestimation for an uncorrected wind of about 20%. This result emphasizes the necessity to take into account the effect of aspect sensitivity in horizontal wind measurements even for a radar with a relatively large oblique beam angle (15°) but with a broad beamwidth (6.8°). Unfortunately, there was no radiosounding close to the radar site able to validate these results, and the wind analyzed by the ECMWF model is, in the jet stream region, weaker than the uncorrected radar wind.

The above results are based on the assumption of horizontal refractive index layers. Larsen and Röttger [1991] have observed persistent refractive index layers tilted from the ground with an angle of about 1° . This tilt angle is comparable to the beam pointing angle offset produced by aspect sensitivity, which for P/P_0 ratios of 10 and 20 dB, are equal to 1.3° and 2.6° , respectively. A more rigorous approach of the horizontal wind correction due to aspect sensitivity needs to take into account the effect of the tilt angle of the atmospheric layers, i.e., to modify the formulation of the effective pointing angle θ_e in (2). As θ_e is relative to the normal of the atmospheric layers, it is a simple matter to correct it, provided the tilt angle and the azimuthal direction of the scattering layer slope are known.

6. Temporal Evolution of Radar Parameters at the Level of the Jet Stream Core

The analysis of aspect sensitivity is now concentrated on a particular level that crosses the most aspect-sensitive region and the core of the upper jet stream. Figure 5 presents a time series of several

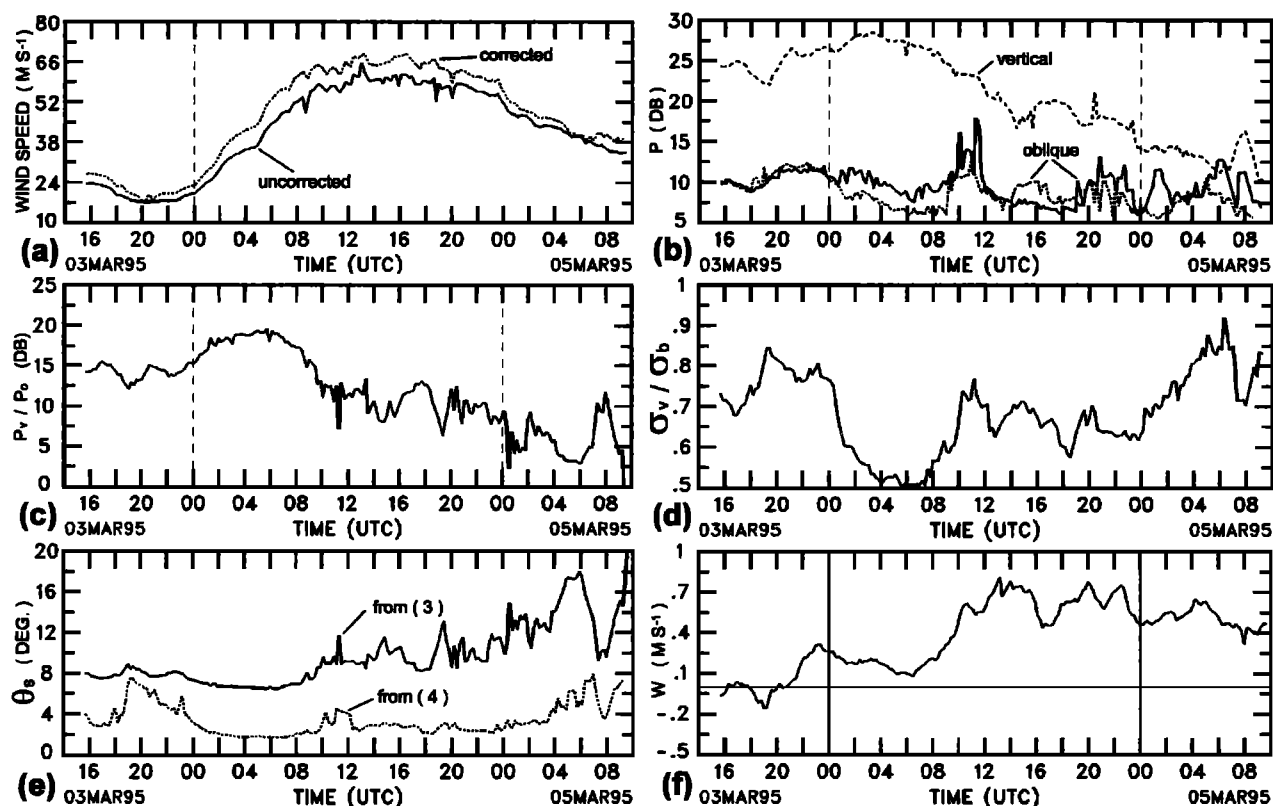


Figure 5. Time series of radar parameters obtained at a height of 9 km: (a) aspect corrected and uncorrected wind speed, (b) signal-to-noise ratio for the three beams corrected for the squared range dependence, (c) aspect ratio P_v/P_o , (d) aspect ratio σ_v/σ_b , (e) aspect angle θ_s derived from equations (3) and (4), and (f) vertical velocity w . A 2-hour running mean was applied, and values deviating more than twice the standard deviation were removed.

radar parameters acquired during the considered period at a height of 9 km. A 2-hour running mean is applied to the data after removal of values deviating more than twice the standard deviation calculated during this 2-hour period.

As shown in Figure 5a, the wind reaches a maximum intensity of about 70 m s^{-1} (aspect corrected) around 1600 UTC on March 4. The uncorrected wind is 8 m s^{-1} weaker. The region of the largest aspect sensitivity values precedes the jet stream core passage and is visible from 0000 to 1000 UTC on March 4. During that period, the maximum aspect sensitivity occurs at 0600 UTC and is characterized in Figures 5c–5e by an aspect ratio P_v/P_o as high as 19 dB and an aspect ratio σ_v/σ_b of 0.5, while the values of the aspect angle θ_s derived from the power and spectral width methods are of 7° and

2° , respectively. Figure 5e clearly shows that both θ_s estimates follow the same general temporal trend, and, as discussed above, the aspect angle value derived from the signal power is always larger by several degrees than the estimate obtained from the vertical spectral width.

Figure 5b presents the signal power observed on the vertical beam and on the two orthogonal oblique beams. The vertical power shows a large variation of about 18 dB during this 42-hour period. On the contrary, the oblique power values oscillate around 10 dB within an interval of about 5 dB. On average, the difference between both oblique powers is close to zero. However, persistent differences, which cannot be attributed to statistical error in the measurements, are clearly apparent in at least two particular regions. The first one lasts 10 hours and is associated with the

largest aspect sensitivity event presented above. The difference between the two oblique signals is about 2 dB and corresponds, considering the orientation of the beams, to aspect sensitivity stronger along the wind than transverse to the wind direction. The reverse is observed during about 5 hours at the time of the passage of the jet stream maximum. Such azimuthal anisotropy of the aspect sensitivity was reported by Yoe *et al.* [1994], Jain *et al.* [1997], and Tsuda *et al.* [1997].

The time series of the vertical velocity w measured with the zenith beam is plotted in Figure 5f. During more than 50% of the considered period w is larger than 0.4 m s^{-1} and reaches a maximum of 0.8 m s^{-1} around the wind maximum. A similar range of values was measured by Yoe *et al.* [1994] during jet stream passages. However, this radar variable has to be taken with caution since several potential sources of error can affect this measurement. A deviation in the vertical pointing angle of 1° can induce, in the presence of horizontal winds of 40 m s^{-1} , a false vertical velocity of 0.7 m s^{-1} which is comparable to the observed values. An equivalent effect arises, in conjunction with aspect sensitivity, from the atmospheric refractivity layer tilted from the horizontal. As mentioned earlier, tilt angles of the order of 1° were measured by Larsen and Röttger [1991] with a VHF interferometric method.

During the passage overhead of the strong aspect sensitivity region, between 0000 and 1000 UTC on March 4, the vertical velocity (Figure 5f) decreases regularly down to a minimum of 0.1 m s^{-1} at 0600 UTC, coincident with the maximum of aspect sensitivity in Figure 5d, and increases afterward. The same temporal trend can be seen in the aspect sensitivity evolution displayed in Figure 5d. The evident similarity in the behavior of these two parameters during a period of 10 hours may not be a simple coincidence, especially when remembering that this period is also associated with a persistent azimuthal anisotropy. The presence of tilted refractivity layers seems a good candidate to explain these observations. The same suggestion was made by Larsen *et al.* [1991] to explain the persistent discrepancy over several hours between the vertical velocity measured from the vertical beam and the one indirectly derived from the combination of radial velocity measurements provided by several oblique beams.

7. Spatial Distribution of the Aspect Sensitivity Fields

During this cut-off low event, the other four VHF radars of the network were in operation. The CRA radar selected for this study, well positioned in the margin of the cut-off low, offered the best and most continuous description of the aspect sensitivity fields. The others radars were either outside of the interesting atmospheric region and/or collected the data during discontinuous short time periods due to instrumental malfunctions. After inspection of this data set, we have selected, for an analysis of the spatial extent of the aspect sensitivity fields, a 10-hour period of common observation of the CRA radar and of the VHF radar located in Toulouse (hereinafter referred to as the TLSE radar). This best comparison period goes from 2200 UTC, March 4 until 0800 UTC, March 5, 1995. As mentioned in section 2, both radars have identical technical characteristics and use the same off-line editing software.

Figure 6 displays, for the selected period, the time-height sections of P_v , P_v/P_o , and σ_v/σ_b provided by the TLSE radar. Time-averaged vertical profiles over this period of these three parameters obtained by the two radars are presented in Figure 7. The radar fields of Figure 6 compare well with those in Figure 3 derived

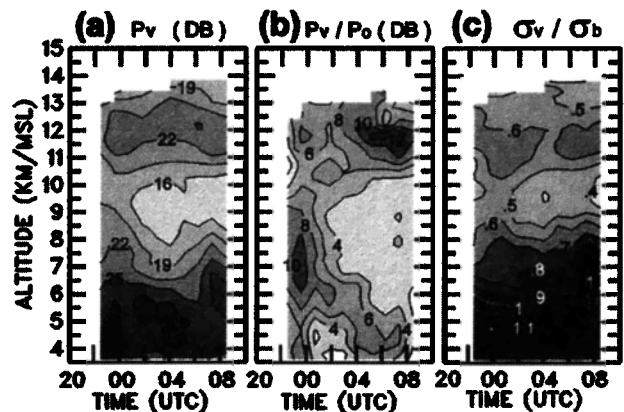


Figure 6. Time-height sections of parameters observed by the VHF radar located in Toulouse during the period from 2200 UTC, March 4 to 0800 UTC, March 5, 1995: (a) vertical echo power P_v (decibels relative to an arbitrary reference power level); (b) vertical to oblique echo power ratio P_v/P_o (decibels); and (c) normalized vertical spectral width σ_v/σ_b . The gray scale is the same as that used in Figure 3.

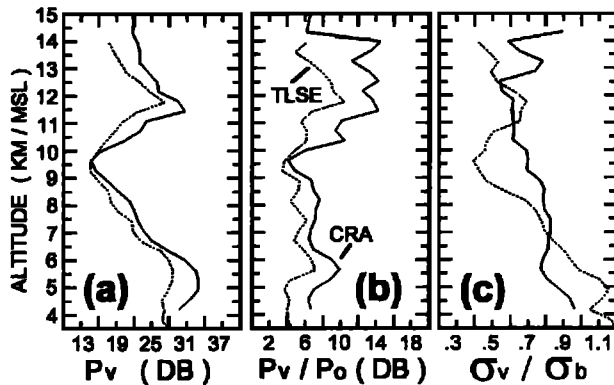


Figure 7. Time-averaged vertical profiles, over the 10-hour period considered in Figure 6, of parameters measured by the CRA radar (solid lines) and the TLSE radar (dotted lines): (a) vertical echo power P_v (decibels relative to an arbitrary reference power level); (b) vertical to oblique echo power ratio P_v/P_o (decibels); and (c) normalized vertical spectral width σ_v/σ_b .

by the CRA radar. The enhancement of the vertical backscattered power P_v in the tropopause region is seen by both radars around 12 km altitude. The vertical evolutions of P_v for these two instruments are also very close in Figure 7a. The descending branch of higher P_v/P_o values observed by the CRA radar (Figure 3b) is also marked in Figure 6b. The TLSE radar has only observed the extremity of this branch, which extends in Figure 6b from an altitude of about 7 km at 2200 UTC on March 4 down to 4 km altitude at about 0600 UTC, March 5. The vertical profiles in Figure 7b show a similar evolution of P_v/P_o with height, but TLSE radar have always observed weaker values of this index of aspect sensitivity. The comparison is not as good for σ_v/σ_b (Figures 3c and 6c). In Figure 7c, there is a relatively large deviation between the measurements by the two radars of this parameter. The TLSE radar presents larger values of σ_v/σ_b at low altitude between 4 and 7 km and smaller values at middle altitude from 7 to 11 km. Since the comparison with P_v/P_o or P_v is relatively satisfactory, the cause of the discrepancy obtained with σ_v/σ_b is unclear. It might result from inaccuracies in the measurement of the vertical spectral width σ_v due, for instance, to the effect of ground clutter imperfectly removed. On the other hand, this parameter is sensitive to small-scale turbulence, and a different turbulence regime at the vertical of the radars might explain part of the observed discrepancies.

When each radar section of Figure 3 is superposed on each corresponding section of Figure 6, the best visual correlation is obtained when TLSE images are shifted by a negative time lag of about 4 hours. A similar result is also obtained with a comparison of the wind field time-height sections. This indicates that the TLSE radar has observed the same atmospheric feature as the CRA radar but with a delay of approximately 4 hours. This comparison made with closely spaced radars brings some insight into the spatial distribution and coherence time of the aspect sensitivity structure, which can be useful in the understanding of the physical process responsible of the anisotropic specular echoes.

8. Conclusion

In this paper, we have examined the aspect sensitivity of VHF radar echoes, through the analysis of the vertical signal enhancement and the narrowing of the vertical spectral width, during a jet stream passage associated with a cut-off low. The goal of this study was to assess the potentiality of a VHF wind profiler to investigate the interactions between stratospheric and tropospheric air. In past studies (several are quoted in the text), large aspect sensitivities have mainly been observed in the lower stratosphere. In the present case study, large aspect sensitivity occurred in about two thirds of the middle and upper troposphere during the 42-hour period analyzed. As shown by the analysis of the ECMWF model, strong aspect sensitivity regions were found to be associated with the tropopause and a tropopause fold, indicating an intrusion of stratospheric air. However, there was a discrepancy between the location of these peculiar atmospheric zones given by the radar and by the numerical model. Particularly, the tropopause fold observed by the radar appears deeper. This result emphasizes the interest in conjointly using numerical model outputs and wind profiler products in order to get a better understanding of the vertical transfers between the stratosphere and the troposphere.

Owing to the aspect sensitivity effect, the underestimation of the horizontal wind speed was found to reach 20% of the nominal value. This points out that even for a radar with a relatively large oblique beam angle (15°) but with a broad beam width (6.8°), aspect sensitivity can be, in certain circumstances, an important source of error that requires to be detected

and corrected. Among the results obtained in this work, which confirm some past observations, is the presence of azimuthal anisotropy over a long period of time and a connection between strong aspect sensitivity and vertical velocity, which might indicate a possible source of error in the measurement of this most important atmospheric parameter. Finally, a comparison of simultaneous observations of two closely spaced VHF radars (100 km) have showed a good spatial coherence of the aspect sensitivity structures.

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- B. Campistron, F. Lohou, and J. P. Pagès, Centre de Recherches Atmosphériques, 65300 Campistrous, France. (e-mail: camb@aero.obs-mip.fr)
- Y. B. Pointin, Laboratoire de Météorologie Physique, Université Blaise Pascal, 24 avenue des Landais, 63177 Aubière, France. (e-mail: pointin@opgc.univ-bpclermont.fr)
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